

11/13/2007

Recall the flow past a sphere. In spherical polar coordinates we have (r, θ, ϕ) and we have axis-symmetric flow so that $\frac{\partial}{\partial \phi} = 0$ and we have u_r, u_θ only, where the flow is moving along the z axis. We expect that

$$u_i = \Phi_{ij}(\vec{x})u_j^{(0)}$$

and

$$p = \frac{1}{a}\mu\Psi(\vec{x})u_j^{(0)}$$

where the force on the sphere is given by

$$F_i = \alpha\mu a u_i^{(0)}$$

for some α . We will use a stream function $\psi(r, \theta)$ where

$$u = \frac{1}{r^2 \sin\theta} \begin{vmatrix} e^{(r)} & r e^{(\theta)} & r \sin\theta e^{(\phi)} \\ \frac{\partial}{\partial r} & \frac{\partial}{\partial \theta} & 0 \\ 0 & 0 & \psi \end{vmatrix}$$

since $u = \nabla \times \left(\frac{\psi}{r \sin\theta} e^{(\phi)}\right)$, where the $\frac{1}{r \sin\theta}$ is to make sure that ψ is constant along the streamlines. Then, we get that

$$u_r = \frac{1}{r^2 \sin\theta} \frac{\partial \psi}{\partial \theta}$$

$$u_\theta = -\frac{1}{r \sin\theta} \frac{\partial \psi}{\partial r}.$$

What are the boundary conditions? We have no normal flow at $r = a$ so that $u_r|_{r=a} = 0$ and no slip so that $u_\theta|_{r=a} = 0$. Hence,

$$\left. \frac{\partial \psi}{\partial r}, \frac{\partial \psi}{\partial \theta} \right|_{r=a} = 0$$

(for all θ on $r = a$). This means that ψ is not varying on the boundary of the sphere so that $\psi = \text{constant}$ on $r = a$ for all θ . We also need the far field condition as $r \rightarrow \infty$. In the far field, we have that

$$u = u_0 e^{(z)}.$$

Then, if we project u in the far field onto $e^{(r)}$ and $e^{(\theta)}$ we get

$$u_r \sim u_0 \cos\theta$$

$$u_\theta \sim -u_0 \sin\theta.$$

Now, using our equations for u in terms of ψ we get that for u_θ

$$-\frac{1}{r \sin\theta} \frac{\partial \psi}{\partial r} \sim -u_0 \sin\theta$$

$$\Rightarrow \frac{\partial \psi}{\partial r} \sim u_0 r \sin^2\theta$$

$$\psi \sim \frac{1}{2} u_0 r^2 \sin^2\theta.$$

This is also the stream function for uniform flow in spherical polars.

This suggests looking for a solution of ψ of the form $f(r) \sin^2\theta$. But before we try this solution, we check to see if this form satisfies the far field condition for u_r .

$$\begin{aligned}\frac{\partial\psi}{\partial\theta} &= u_0 r^2 \sin\theta \\ \frac{1}{r^2 \sin\theta} \frac{\partial\psi}{\partial\theta} &= u_0 \cos\theta.\end{aligned}$$

So we expect that $\psi \sim \frac{1}{2} u_0 r^2 \sin^2\theta$ in the far field. Remember that for Stokes flows we have

$$\nabla^2 \omega = 0$$

since we have that $0 = -\nabla p + \mu \nabla^2 u$ and taking the curl gives us the above relationship. How do we write Laplace's equation in spherical polar coordinates? We have that

$$\begin{aligned}u &= \nabla \times \left(\frac{\psi}{r \sin\theta} e^{(\phi)} \right) \\ \omega &= \nabla \times u\end{aligned}$$

which gives us

$$\omega = \frac{1}{r^2 \sin\theta} \begin{vmatrix} e^{(r)} & r e^{(\theta)} & r \sin\theta e^{(\phi)} \\ \frac{\partial}{\partial r} & \frac{\partial}{\partial \theta} & 0 \\ u_r & r u_\theta & 0 \end{vmatrix}.$$

The only nonzero vorticity is in the $e^{(\phi)}$ direction (i.e. $\omega = \omega_\phi e^{(\phi)}$). Hence,

$$\begin{aligned}\omega_\phi &= \frac{1}{r} \left[\frac{\partial}{\partial r} (r u_\theta) - \frac{\partial u_r}{\partial \theta} \right] \\ &= \frac{1}{r} \left[\frac{\partial}{\partial r} \left(-\frac{1}{\sin\theta} \frac{\partial\psi}{\partial r} \right) - \frac{\partial}{\partial \theta} \left(\frac{1}{r^2 \sin\theta} \frac{\partial\psi}{\partial \theta} \right) \right] \\ &= -\frac{1}{r \sin\theta} \left[\frac{\partial^2 \psi}{\partial r^2} + \frac{\sin\theta}{r^2} \frac{\partial}{\partial \theta} \left(\frac{1}{\sin\theta} \frac{\partial\psi}{\partial \theta} \right) \right].\end{aligned}$$

That is,

$$\omega_\phi = -\frac{1}{r \sin\theta} \left[\frac{\partial^2 \psi}{\partial r^2} + \frac{\sin\theta}{r^2} \frac{\partial}{\partial \theta} \left(\frac{1}{\sin\theta} \frac{\partial\psi}{\partial \theta} \right) \right].$$

Now we have similarity between ω_ϕ and u . Define $D^2 \psi = \frac{\partial^2 \psi}{\partial r^2} + \frac{\sin\theta}{r^2} \frac{\partial}{\partial \theta} \left(\frac{1}{\sin\theta} \frac{\partial\psi}{\partial \theta} \right)$. Then,

$$\omega = -\frac{1}{r \sin\theta} D^2 \psi e^{(\phi)}.$$

So we have that

$$\begin{aligned}\omega &= \nabla \times \left(\nabla \times \left\{ \frac{\psi}{r \sin\theta} e^{(\phi)} \right\} \right) \\ &= -\frac{1}{r \sin\theta} D^2 \psi e^{(\phi)}.\end{aligned}$$

We still need to find $\nabla^2 \omega$. Let us take

$$\nabla \times (\nabla \times \omega) = \nabla(\nabla \cdot \omega) - \nabla^2 \omega$$

and note that $\nabla \cdot \omega = \nabla \cdot (\nabla \times u) \equiv 0$ so that

$$\nabla \times (\nabla \times \omega) = -\nabla^2 \omega.$$

So we want to show that $\nabla \times (\nabla \times \omega) = 0$. ω is only in the $e^{(\phi)}$ direction. If we replace ψ by the term $(-D^2\psi)$ we get

$$\begin{aligned} \nabla \times (\nabla \times \omega) &= -\nabla \times \left(\nabla \times \frac{D^2\psi}{r \sin\theta} e^{(\phi)} \right) \\ &= \frac{1}{r \sin\theta} D^2(D^2\psi) e^{(\phi)}. \end{aligned}$$

Hence, we solve

$$D^2(D^2\psi) = 0.$$

This is now a scalar PDE. First, how do we solve $D^2g = 0$? Let us consider $g(r, \theta) = G(r) \times f(\theta)$. We know that the sphere is isotropic so it not angularly dependent. However, in the far field ψ is angularly dependent. So let us try $\psi = f(r)\sin^2\theta$ since this is the form of ψ in the far field. Then,

$$\begin{aligned} D^2\psi &= \frac{d^2f}{dr^2} \sin^2\theta + \frac{\sin\theta}{r^2} f(r) \frac{d}{d\theta} \left(\frac{1}{\sin\theta} 2 \sin\theta \cos\theta \right) \\ &= \frac{d^2f}{dr^2} \sin^2\theta - \frac{2\sin^2\theta}{r^2} f(r) \\ &= \left\{ \frac{d^2}{dr^2} - \frac{2}{r^2} \right\} f(r) \sin^2\theta. \end{aligned}$$

So we set $F(r) = \left(\frac{d^2}{dr^2} - \frac{2}{r^2} \right) f(r)$ so that

$$\begin{aligned} D^2(D^2\psi) &= \left(\frac{d^2}{dr^2} - \frac{2}{r^2} \right) F(r) \sin\theta \\ &= 0. \end{aligned}$$

This is a second order ODE for $F(r)$. That is, we need to solve

$$F''(r) - \frac{2}{r^2} F(r) = 0.$$

Note that this is a homogeneous equation with the same dimensionality for F'' and F/r^2 . Hence, $F \propto r^\alpha$. Then, $F'' \propto \alpha^2 r^{\alpha-2}$ and

$$\begin{aligned} \alpha(\alpha-1)r^{\alpha-2} - 2r^{\alpha-2} &= 0 \\ \alpha^2 - \alpha - 2 &= 0 \\ (\alpha-2)(\alpha+1) &= 0. \end{aligned}$$

We get that $\alpha = 2, -1$. Then,

$$F(r) \propto r^2, r^{-1}.$$

Next, we find $\left(\frac{d^2}{dr^2} - \frac{2}{r^2} \right) f = Pr^2 + Qr^{-1}$. We can try $f \propto r^\alpha$. Again, we get that

$$(\alpha^2 - \alpha - 2)r^{\alpha-2} \propto r^2, r^{-1}.$$

If we set RHS = 0 then we get $\alpha = 2, -1$. If we want an r^2 term, then $\alpha = 4$ and if we want r^{-1} we need $\alpha = 1$. This is the method of undetermined coefficients (we guess at the solution from the RHS). We look for the solution of the form

$$f(r) = Ar^4 + Br^2 + Cr + Dr^{-1}.$$

We get that for Ar^4 we have $A(12r^2 - 2r^2) = Pr^2$ so that $10A = P$. Also we have that Cr becomes $C(-2r^{-1}) = Qr^{-1}$ so that $-2C = Q$. So we have that

$$f(r) = Ar^4 + Br^2 + Cr + Dr^{-1},$$

and $\psi = f(r) \sin^2\theta$ and $\psi \sim \frac{1}{2}r^2 \sin^2\theta u_0$ as $r \rightarrow \infty$. As $r \rightarrow \infty$ A must be identically zero. Matching the r^2 term, we also get that $B = \frac{1}{2}u_0$. Then,

$$f(r) = \frac{1}{2}u_0r^2 + Cr + \frac{D}{r}.$$

For $r = a$, remember that $\psi = \text{constant}$ for all θ and $\left. \frac{\partial \psi}{\partial r} \right|_{r=a} = 0$. Hence, $f(a) = 0$ and $\frac{df}{dr} = 0$ on $r = a$ otherwise we have a \sin^2 variation for ψ if we don't let $f(a) = 0$. Then

$$\begin{aligned} 0 &= \frac{1}{2}u_0a^2 + Ca + \frac{D}{a} \\ \Rightarrow 0 &= \frac{1}{2}u_0a + C + \frac{D}{a^2}. \end{aligned}$$

After some algebra, we get that $C = -\frac{3}{4}au_0$ and $D = \frac{1}{4}u_0a^3$. Then,

$$\psi(r, \theta) = \frac{1}{2}u_0r^2 \sin^2\theta \left\{ 1 - \frac{3a}{2r} + \frac{1}{2} \frac{a^3}{r^3} \right\}.$$

Note that $Pr^2 + Qr^{-1}$ come from the vorticity so that A, C are linked with the vorticity. Whereas B and D have no vorticity (irrotational or potential flows) (but C has vorticity).

What is the vorticity? We showed that

$$\begin{aligned} \omega_\phi &= -\frac{1}{r \sin\theta} D^2 \psi, \\ \psi &= f(r) \sin^2\theta \\ \Rightarrow \omega_\phi &= -\frac{1}{r \sin\theta} \left\{ \frac{d^2}{dr^2} - \frac{2}{r^2} \right\} f(r) \sin^2\theta \\ &= -\frac{1}{r} \left\{ \frac{d^2}{dr^2} - \frac{2}{r^2} \right\} f(r) \sin\theta \\ &= \frac{2C \sin\theta}{r^2}. \end{aligned}$$

This peaks at $\theta = \frac{\pi}{2}$ so that $\omega_\phi = \frac{2C}{r^2} = -\frac{3}{2}u_0 \frac{a}{r^2}$. Hence, the maximum vorticity is at $\theta = \pi/2, r = a$ and we get $\|\omega_\phi\|_\infty = -\frac{3}{2}u_0 \frac{1}{a}$.

As for the pressure term, we have that

$$\begin{aligned} \nabla p &= \mu \nabla^2 u \\ &= -\mu (\nabla \times \omega) \end{aligned}$$

since $\nabla \cdot u = 0$ because $u = \nabla \times \psi$ and the flow is incompressible. That is, we are using

$$\nabla \times (\nabla \times u) = \nabla(\nabla \cdot u) - \nabla^2 u.$$

So the pressure is determined by the C term. One can find

$$\frac{1}{r} \frac{\partial p}{\partial \theta} = \frac{\mu}{r \sin \theta} \frac{\partial}{\partial r} (r \sin \theta \omega_\phi)$$

and so

$$\begin{aligned} p &= p_0 + 2 \frac{\mu C}{r^2} \cos \theta \\ &= p_0 - \frac{3}{2} \mu u_0 \frac{a}{r^2} \cos \theta, \end{aligned}$$

where p_0 is the ambient pressure. The r dependence on p vanishes because ω is only in the $\hat{\phi}$ direction. So on the $\theta = \pi$ axis, we have a higher pressure as compared with $\theta = 0$. So we have a net pressure force. Let us look at the pressure term a little more closely.

$$p - p_0 = -\frac{3}{2} \mu u_0 \frac{a}{r^2} \cos \theta.$$

Recall that $u^{(0)} = u_0 e^{(z)}$. If we take $x \cdot u^{(0)} = u_0 r \cos \theta$. Then, the above equation for the pressure becomes

$$p - p_0 = -\frac{3}{2} \mu a \frac{x \cdot u^{(0)}}{r^3}.$$

Now note that $\frac{x \cdot u^{(0)}}{r^3}$ is a particular solution to Laplace's equation. Note that we have $\nabla^2 p = 0$ and we have a potential dipole term (recall that $\nabla p = -\mu(\nabla \times \omega)$ and if take the divergence of both sides we get $\nabla^2 p = 0$).

The pressure force term is

$$\begin{aligned} F_z^{(p)} &= - \oint_S p n \cdot e^{(z)} dS \\ &= \frac{3}{2} \mu u_0 \oint_S \frac{a}{r^2} \cos^2 \theta dS \\ &= \frac{3}{2} \mu u_0 \frac{1}{a} 2\pi a^2 \int_0^\pi \cos^2 \theta \sin \theta d\theta \\ &= \frac{3}{2} \mu u_0 \frac{1}{a} 2\pi a^2 \left(-\frac{1}{3} \cos^3 \theta \right) \end{aligned}$$

By symmetry we know that the force must be in the streamwise direction. Thus,

$$F_z^{(p)} = 2\mu\pi a u_0.$$

We also have that

$$F_i = \oint_S \sigma_{ij} n_j dS$$

and $n = e^{(r)}$ so we need $\sigma_{rr} e^{(r)} + \sigma_{\theta r} e^{(\theta)}$. Then,

$$\begin{aligned} \sigma_{rr} &= -p + 2\mu \frac{\partial u_r}{\partial r} \\ \sigma_{\theta r} &= \mu \left(\frac{\partial u_\theta}{\partial r} - \frac{u_\theta}{r} + \frac{1}{r} \frac{\partial u_r}{\partial \theta} \right). \end{aligned}$$

Finally, we get that

$$F_z^{(v)} = 4\pi\mu a u_0.$$

Hence, the total force is

$$F = 6\pi a \mu u_0 e^{(z)}.$$

The B and D terms do not contribute to the force. Only the C term matters. The B term corresponds to uniform flow so we did not expect that to contribute a force. As for the D term, recall that $\psi \sim D/r$ so that $u \sim D/r^3$ and the stresses will look like D/r^4 . So the stresses vanish like $1/r^4$ while the surface area increases as r^2 so these terms vanish.

11/15/2007

Multipole Methods:

Suppose we have a sphere of radius a moving with velocity v centered at the origin. Then,

$$\begin{aligned} u_i(x) &= \frac{3}{4}a \left[\frac{\delta_{ij}}{r} + \frac{x_i x_j}{r^3} \right] v_j + \frac{1}{4}a^3 \left[\frac{\delta_{ij}}{r^3} - 3 \frac{x_i x_j}{r^5} \right] v_j \\ p(x) &= \frac{3}{2}a\mu \frac{v \cdot x}{r^3}. \end{aligned}$$

The nice thing here is that we have the formula in terms of regular cartesian coordinates. For the u_i term we call the first term on the RHS as the Stokeslet and the second term as the Potential flow-dipole. Recall that

$$\psi = \frac{1}{2}u_0 r^2 \sin^2 \theta \left\{ 1 - \frac{3a}{2r} + \frac{1}{2} \frac{a^3}{r^3} \right\},$$

where the last term $\frac{1}{2} \frac{a^3}{r^3}$ has a zero vorticity. Similarly, $\frac{\delta_{ij}}{r^3} - 3 \frac{x_i x_j}{r^5}$ has zero vorticity. We have that

$$\begin{aligned} u_i^{(2)} &= \left[\frac{\delta_{ij}}{r^3} - 3 \frac{x_i x_j}{r^5} \right] B_j \\ &= \frac{\partial}{\partial x_i} \left(\frac{x_j B_j}{r^3} \right) \end{aligned}$$

and we have that

$$u^{(2)} = \nabla \left(\frac{x \cdot B}{r^3} \right)$$

and we have a potential flow: $\nabla \times u^{(2)}$ (the superscript only refers to the respective term and does not denote a derivative). For the first term,

$$u_i^{(1)} = \left[\frac{\delta_{ij}}{r} + \frac{x_i x_j}{r^3} \right] A_j.$$

Then,

$$\begin{aligned}
\omega_i &= \epsilon_{ijk} \frac{\partial}{\partial x_j} u_k^{(1)} \\
&= \epsilon_{ijk} \frac{\partial}{\partial x_j} \left[\frac{\delta_{km}}{r} + \frac{x_k x_m}{r^3} \right] A_m \\
&= \epsilon_{ijk} \left\{ \delta_{km} \left(-\frac{1}{r^2} \right) \frac{x_j}{r} + \frac{\delta_{kj} x_m}{r^3} + \frac{\delta_{mj} x_k}{r^3} - 3 \frac{x_k x_m x_j}{r^5} \right\} A_m.
\end{aligned}$$

Note that the last term is symmetric in j, k and ϵ_{ijk} is antisymmetric in j, k so this term vanishes after summation. Similarly, with $\frac{\delta_{kj} x_m}{r^3}$. Hence,

$$\begin{aligned}
\omega_i &= \epsilon_{ijk} \left\{ \delta_{km} \left(-\frac{1}{r^2} \right) \frac{x_j}{r} + \frac{\delta_{mj} x_k}{r^3} \right\} A_m \\
&= 2\epsilon_{ijk} \left(\frac{A_j x_k}{r^3} \right),
\end{aligned}$$

by switching j, k is the latter term. Furthermore,

$$\begin{aligned}
\omega &= 2 \frac{A \times x}{r^3}. \\
A &= \frac{3}{4} a \vec{V} \\
&= \frac{3}{4} a V e^{(z)}
\end{aligned}$$

and

$$e^{(z)} \times x = e^{(\phi)} r \sin\theta$$

so that

$$\omega_\phi = \frac{3}{2} \frac{a V \sin\theta}{r^3}.$$

We also need to verify that $\nabla \cdot u^{(1)} = 0$ ($r \neq 0$). Last time we showed that the fluid force on the sphere is given by

$$\vec{F} = -6\pi a \mu \vec{V}.$$

On the other hand, the force of the sphere on the fluid is

$$\begin{aligned}
\hat{\vec{F}} &= -\vec{F} \\
&= 6\pi a \mu \vec{V}.
\end{aligned}$$

Comparing this to $A = \frac{3}{4} a \vec{V}$, we get that

$$\vec{A} = \frac{\hat{\vec{F}}}{8\pi\mu}.$$

As for the pressure term,

$$p = \frac{1}{4\pi} \frac{\hat{F} \cdot x}{r^3}.$$

All of this comes from the Stokeslet term, which we can write as

$$\begin{aligned} u &= \frac{1}{8\pi\mu} \left\{ \frac{\delta_{ij}}{r} + \frac{x_i x_j}{r^3} \right\} F_j \\ p &= \frac{F_j x_j}{4\pi r^3}. \end{aligned}$$

One can verify that

$$\begin{aligned} 0 &= -\nabla p + \mu \nabla^2 u + F \delta(x) \\ 0 &= \nabla \cdot u. \end{aligned}$$

So if we take any sphere, radius R , and evaluate the force on the fluid, it always equals F . Note that $F\delta(x)$ is merely notation.

Boundary Conditions:

As $r \rightarrow \infty$, $u \rightarrow 0$. At $r = a$, we have

$$\begin{aligned} u_i(x) &= \frac{3}{4} \left[\delta_{ij} + \frac{x_i x_j}{a^2} \right] v_j + \frac{1}{4} \left[\delta_{ij} - 3 \frac{x_i x_j}{a^2} \right] v_j \\ &= v_i, \end{aligned}$$

so that we satisfy the no slip boundary conditions.

If we are dealing with a clean gas bubble with radius a we have that

$$u \cdot n = v \cdot n, \text{ on } r = a.$$

or $(v - u) \cdot n = 0$. Also, the shear stress is $\cong 0$ and so

$$\begin{aligned} \epsilon_{kmi} n_m (\sigma_{ij} n_j) &= 0 \\ n \times (\sigma \cdot n) &= 0. \end{aligned}$$

Then,

$$u_i(x) = \left[\frac{\delta_{ij}}{r} + \frac{x_i x_j}{r^3} \right] A_j + \left[\frac{\delta_{ij}}{r^3} - 3 \frac{x_i x_j}{r^5} \right] B_j,$$

for some appropriate A, B . After some calculation, we find that

$$\begin{aligned} B &= 0, \\ A &= \frac{4\pi a \mu V}{8\pi \mu} \\ &= \frac{1}{2} a V. \end{aligned}$$

For other example: liquid sphere, external ($r > a$) flow is same format. In general, the force multipole expression can be written as

$$0 = -\frac{\partial p}{\partial x_i} + \mu \nabla^2 u_i + \left\{ F_i^{(0)} \delta(x) + F_{ij}^{(1)} \frac{\partial}{\partial x_j} \delta(x) + F_{ijk}^{(2)} \frac{\partial^2}{\partial x_j \partial x_k} \delta(x) + \dots \right\}.$$

Can we always neglect fluid inertia? In Stokes flow, $u \propto \frac{Va}{r}$. This is when we have a moving sphere, where the fluid is at rest at ∞ . Recall that our equation of motion is

$$\rho \frac{Du_i}{Dt} = - \frac{\partial p}{\partial x_i} + \mu \nabla^2 u_i$$

and we have that $\text{Re} = \frac{Va}{\nu} \ll 1$ so we can ignore the inertia term. Does this hold away from the sphere? The inertia term is $\rho u \cdot \nabla u$. Some quick dimensional analysis gives us

$$\begin{aligned} \rho u \cdot \nabla u &\rightarrow \rho \frac{V^2 a^2}{r^2} \frac{1}{r} \\ \mu \nabla^2 u &\rightarrow \mu \frac{Va}{r} \frac{1}{r^2}. \end{aligned}$$

This gives us that the ratio of Va to μ/p is $\frac{Va}{\nu}:1$.

A flow past a fixed sphere, we have that $u = u^{(0)}$ as $r \rightarrow \infty$ and we have that u behaves like $u^{(0)} - \frac{u^{(0)}a}{r}$. Hence, the inertial term is given by

$$u \cdot \nabla u \sim u^{(0)} \cdot \frac{u^{(0)}a}{r^2}$$

for $r \gg a$. Then,

$$\nu \nabla^2 u \sim \nu \frac{u^{(0)}a}{r^3}.$$

If we look at the ratio of terms, we get $\frac{u_0 r}{\nu}:1$. For a fixed sphere, we require that $\frac{u_0 r}{\nu} \sim \mathcal{O}(1)$ eventually. So we should correct for this because we need this to be small. Hence, we introduce the Oseen correction to drag force:

$$F = 6\pi a \mu \vec{u} \left\{ 1 + \frac{3}{8} \text{Re} \right\}.$$

In 3D, consistent stokes solutions are possible. In 2D, flow past a cylinder is a problem.

Hele – Shaw Cell:

We have a flow between glass plates, narrow gap h . For the length scales, we have h, D, L where $h \ll D \ll L$. In a steady flow, we have

$$\begin{aligned} u \cdot \nabla u &= - \frac{1}{\rho} \nabla p + \nu \nabla^2 u \\ \nabla \cdot u &= 0 \end{aligned}$$

where

$$\mu \nabla^2 u = \nu \left\{ \frac{u_0}{D^2} + \frac{u_0}{D^2} + \frac{u_0}{h^2} \right\}$$

where the RHS refers to x_1, x_2 and x_3 terms. For u_1, u_2 we have $u_0 \frac{u_0}{D}$. The nonlinear terms versus the viscous terms are

$$\begin{aligned} \frac{u_0^2}{D} &\text{ vs } \nu \frac{u_0}{h^2} \\ \downarrow & \\ \frac{u_0 h}{\nu} \frac{h}{D} &: 1 \end{aligned}$$

where $\frac{u_0 h}{\nu}$ is small and $\frac{h}{D}$ is very small. Hence the effect of the nonlinear terms are negligible. So, for u_1, u_2 we have that

$$0 = -\frac{1}{\rho} \nabla p + \nu \nabla^2 u.$$

Hence, we would get a parabolic velocity profile for u_1, u_2 in x_3 and $u_1, u_2 \propto \nabla p$. Also,

$$0 = \frac{\partial p}{\partial x_1} + \mu \frac{\partial^2 u_1}{\partial x_3^2}$$

and $u_1 = 0$ at $x_3 = 0, h$. Also, $\frac{\partial}{\partial x_3} \left(\frac{\partial p}{\partial x_1} \right)$ is negligible. In summary,

$$\begin{aligned} (u_1, u_2) &\propto \left(-\frac{\partial p}{\partial x_1}, -\frac{\partial p}{\partial x_2} \right) \\ \nabla \cdot u &= 0 \\ \nabla^2 p &= 0. \end{aligned}$$

So we actually observe 2D potential flow past a circle.